

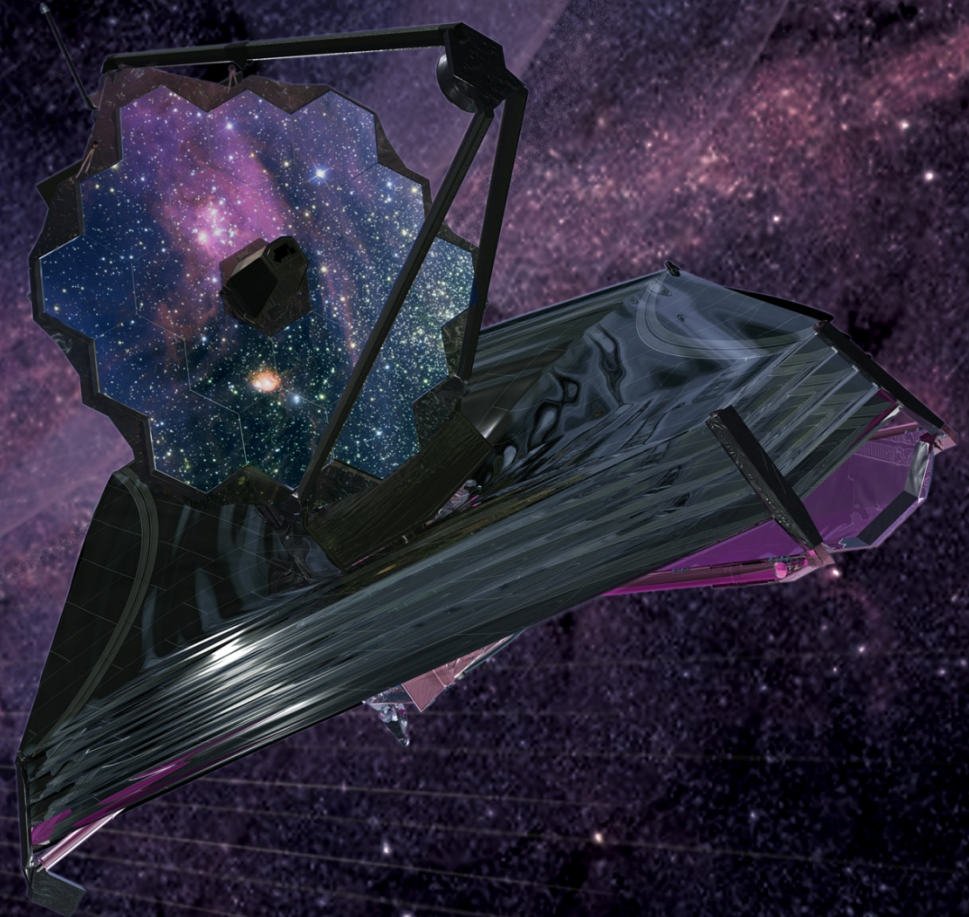
# Status of the JWST Science Instruments

**Matt Greenhouse**

JWST Project Office

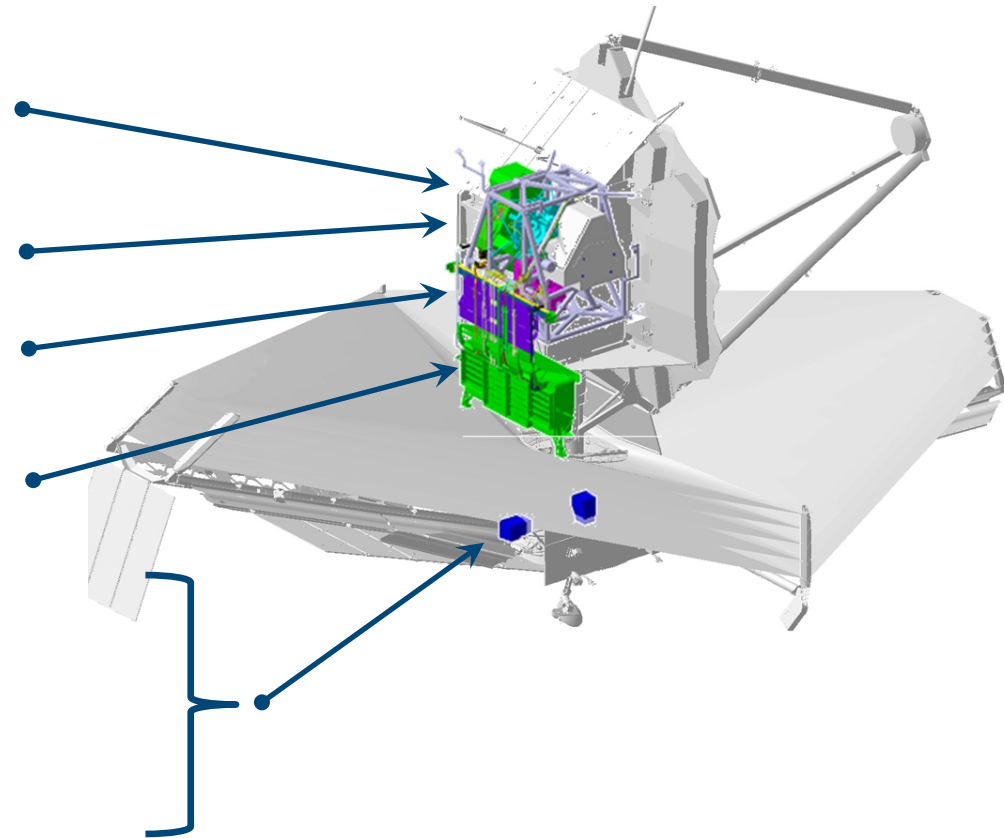
NASA Goddard Space Flight Center

9 January 2012

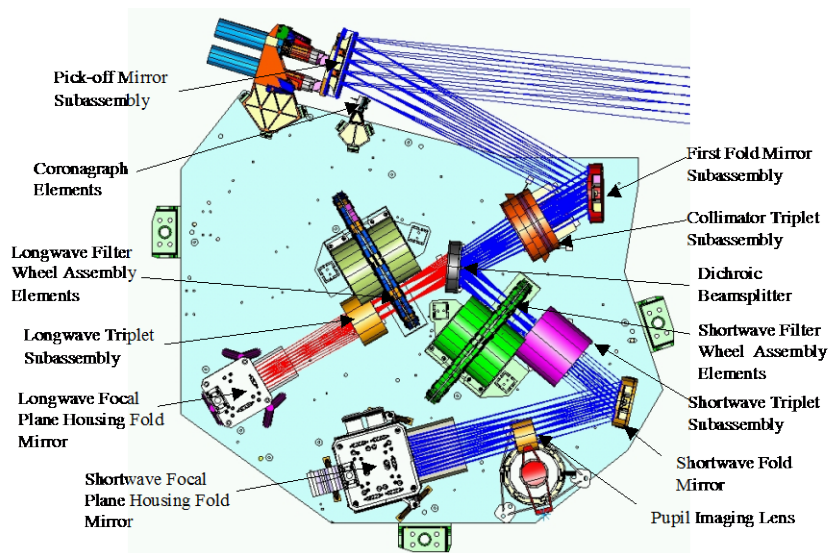
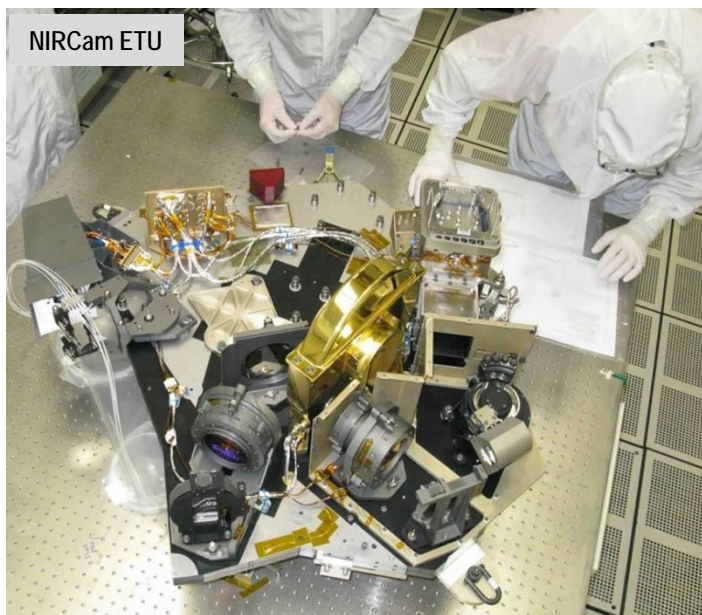


# The Integrated Science Instrument Module (ISIM) is the science instrument payload of the JWST

- ISIM is one of three elements that together make up the JWST space vehicle
  - Approximately 1.4 metric tons, ~20% of JWST by mass
  - Completed its Critical Design Review during 2009 and is currently in integration and test
- The ISIM system consists of:
  - Four science instruments
    - NIRCam, NIRSpec, MIRI, FGS
  - Nine instrument support systems:
    - Optical metering structure system
    - Electrical Harness System
    - Harness Radiator System
    - ISIM electronics compartment
    - ISIM Remote Services Unit
    - Cryogenic Thermal Control System
    - Command and Data Handling System
    - Flight Software System
    - Operations Scripts System



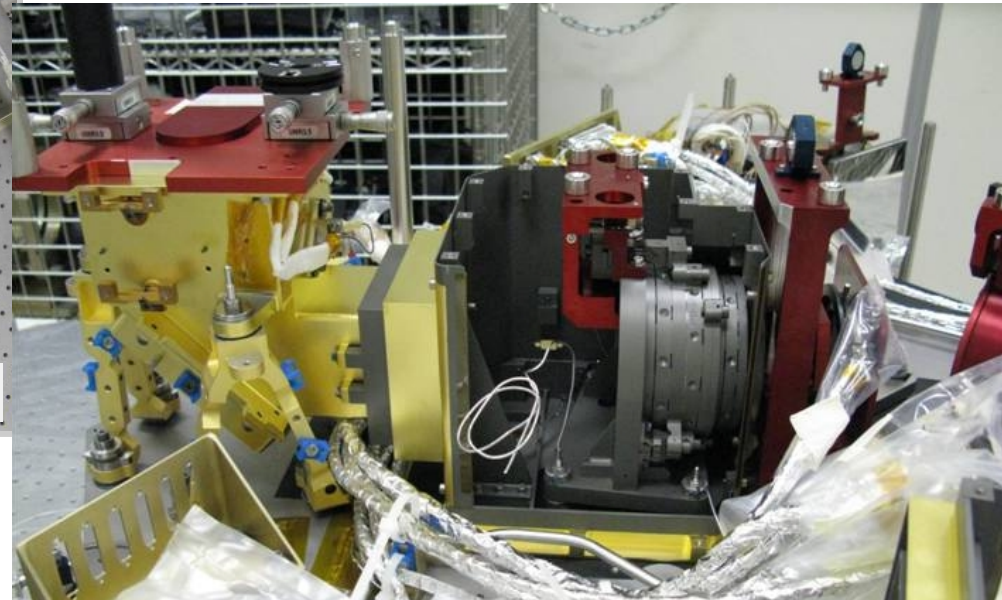
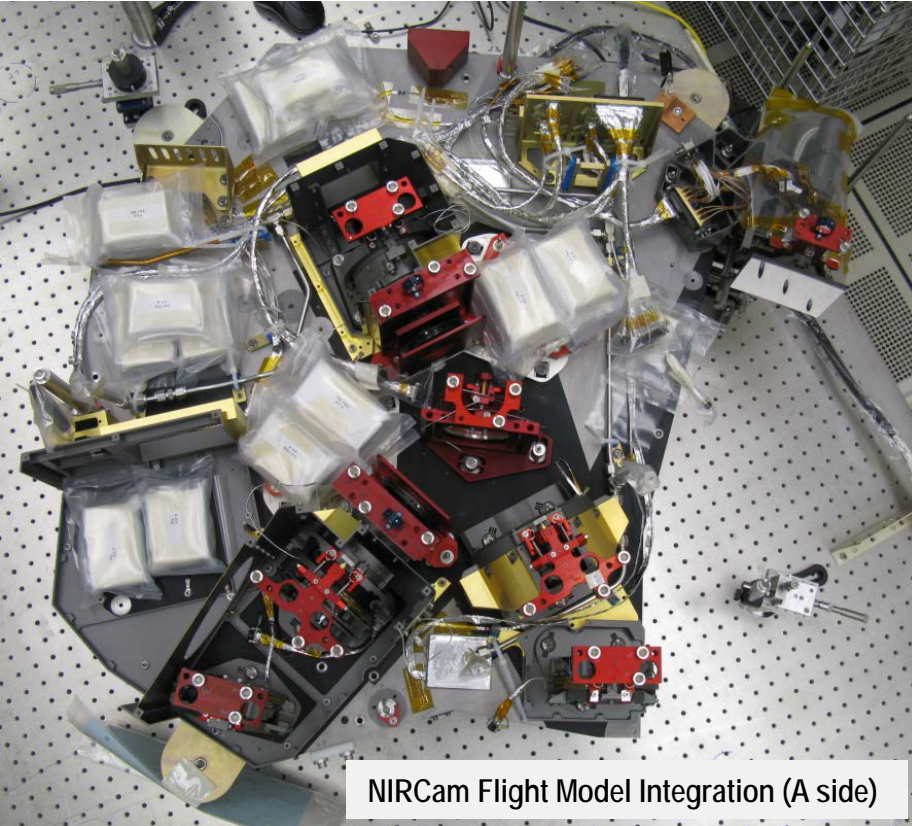
# NIRCam will provide the deepest near-infrared images ever and will identify primeval galaxy targets for the NIRSpec



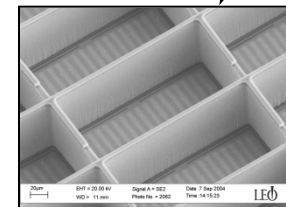
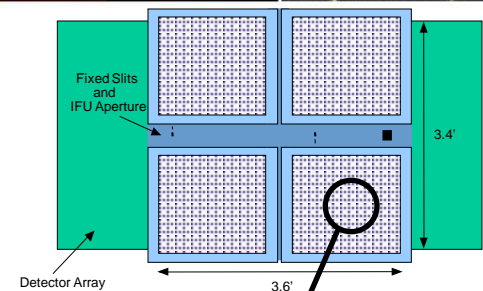
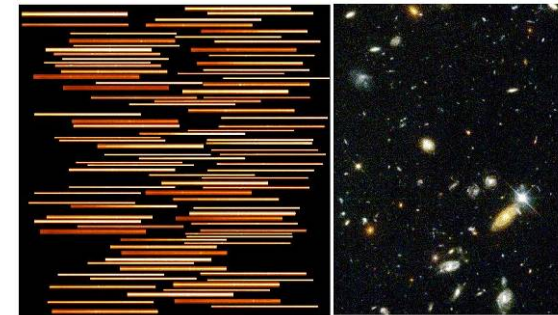
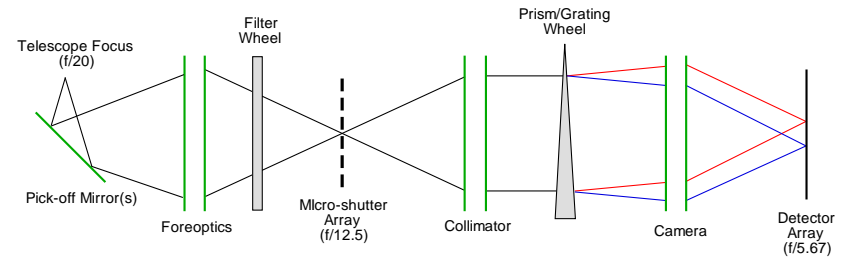
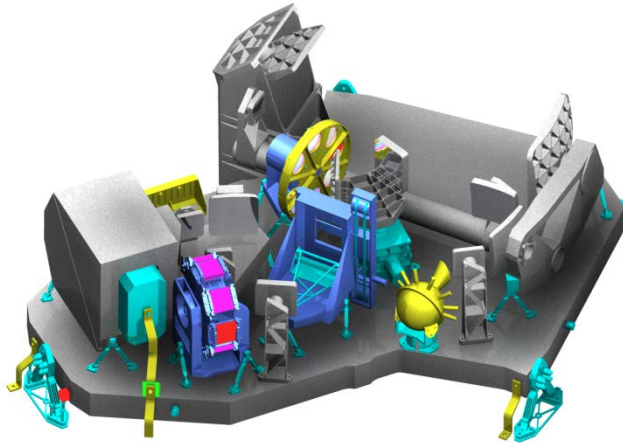
- Developed by the University of Arizona with Lockheed Martin ATC
  - Operating wavelength: 0.6 – 5.0 microns
  - Spectral resolution: 4, 10, 100 (filters + grism), coronagraph
  - Field of view: 2.2 x 4.4 arc minutes
  - Angular resolution (1 pixel): 32 mas < 2.3 microns, 65 mas > 2.4 microns
  - Detector type: HgCdTe, 2048 x 2048 pixel format, 10 detectors, 40 K passive cooling
  - Refractive optics, Beryllium structure
- Supports telescope wavefront sensing

# NIRCam is on schedule for delivery during 2012

Flight model cryo-vacuum testing begins during March



# The NIRSpec will acquire near-infrared spectra of up to 100 objects in a single exposure



- Developed by the European Space Technology Center (ESTEC) with Astrium and Goddard Space Flight Center

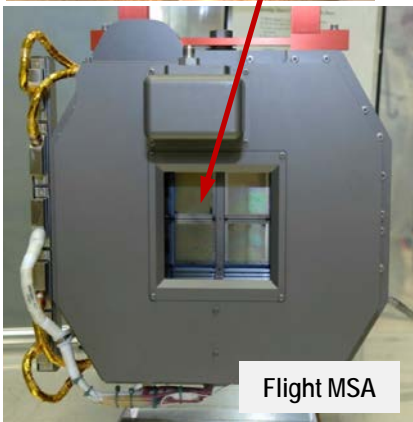
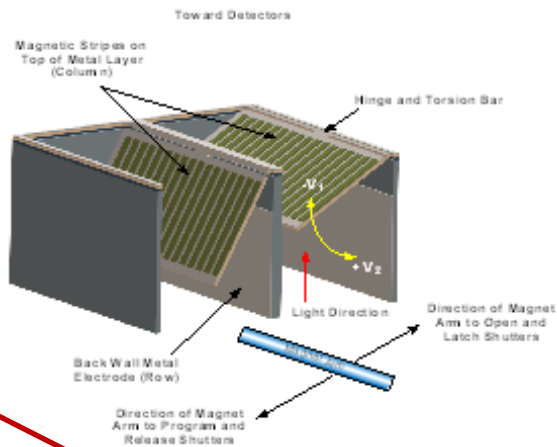
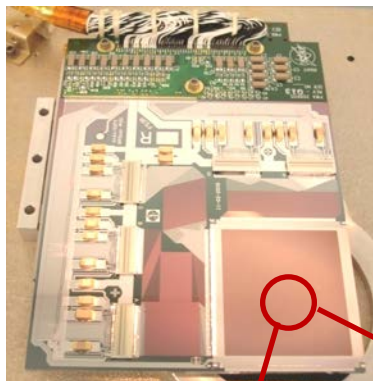
- Operating wavelength: 0.6 – 5.0 microns
- Spectral resolution: 100, 1000, 3000
- Field of view: 3.4 x 3.4 arc minutes
  - Aperture control:

- Programmable micro-shutters, 250,000 pixels
- Fixed long slits & transit spectroscopy aperture
- Image slicer (IFU) 3x3 arc sec

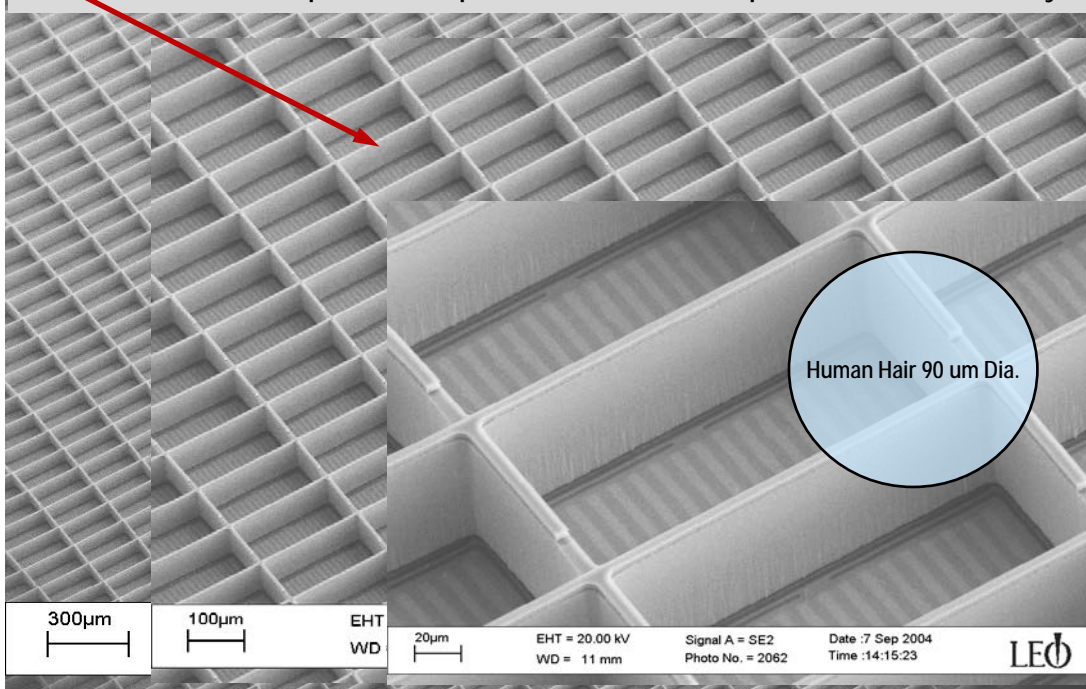
- Detector type: HgCdTe, 2048 x 2048 format, 2 detectors, 37 K passive cooling
- Reflective optics, Silicon Carbide structure and optics

# Aperture control: 250,000 programmable micro-shutters

## System flight qualified and delivered to ESA June 2010



203 x 463 mas shutter pixel clear aperture, 267 x 528 mas pitch, 4 x 171 x 365 array



# NIRSpec delivery expected during early 2013



# The MIRI instrument will characterize circumstellar debris disks, extra-solar planets, and the evolutionary state of high redshift galaxies

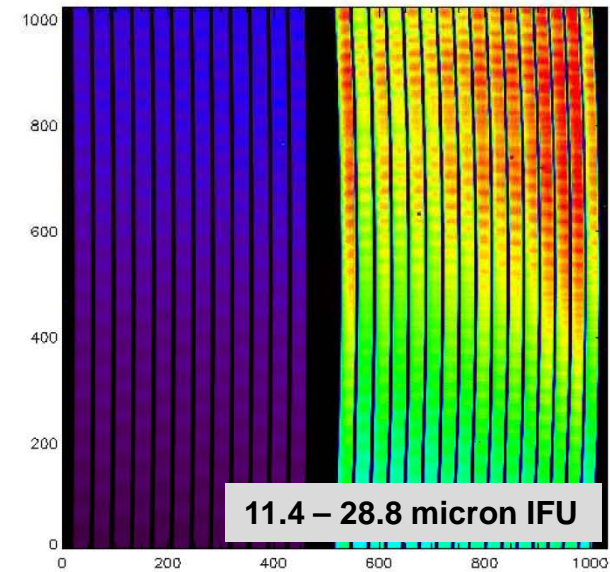
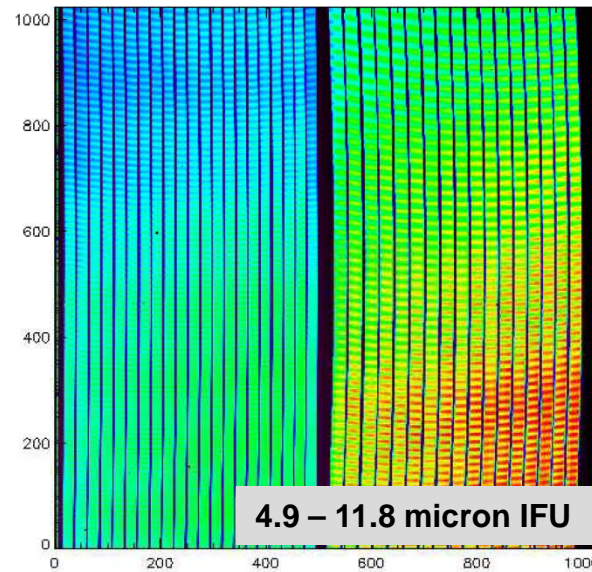
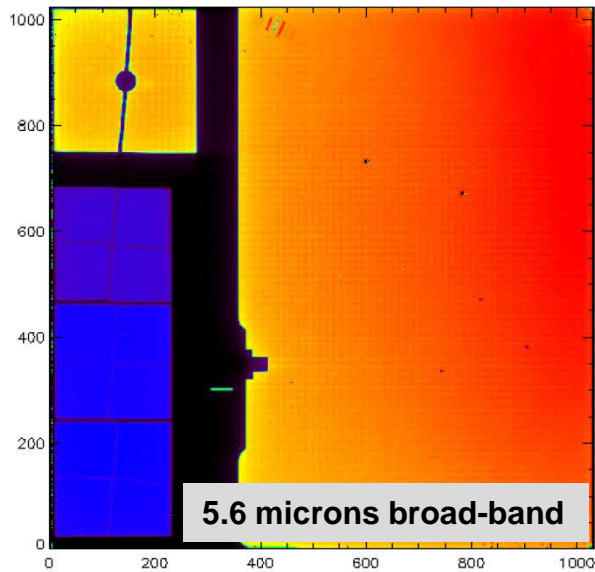
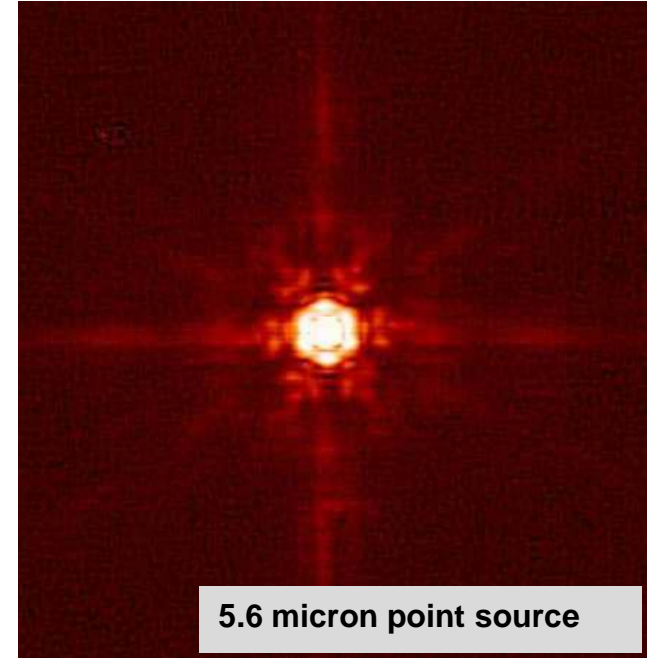
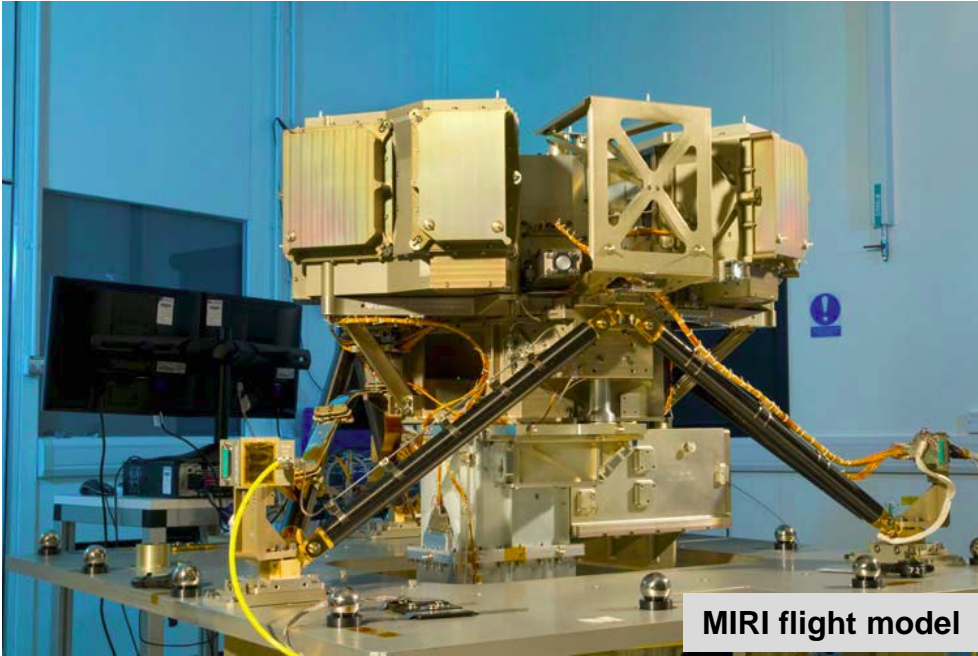


- Developed by a consortium of 10 European countries and NASA/JPL
  - Operating wavelength: 5 - 29 microns
  - Spectral resolution: 5, 100, 2000
  - Broad-band imagery: 1.9 x 1.4 arc minutes FOV
  - Coronagraphic imagery
  - Spectroscopy:
    - R100 long slit spectroscopy 5 x 0.2 arc sec
    - R2000 spectroscopy 3.5 x 3.5 and 7 x 7 arc sec FOV integral field units
  - Detector type: Si:As, 1024 x 1024 pixel format, 3 detectors, 7 K cryo-cooler
  - Reflective optics, Aluminum structure and optics

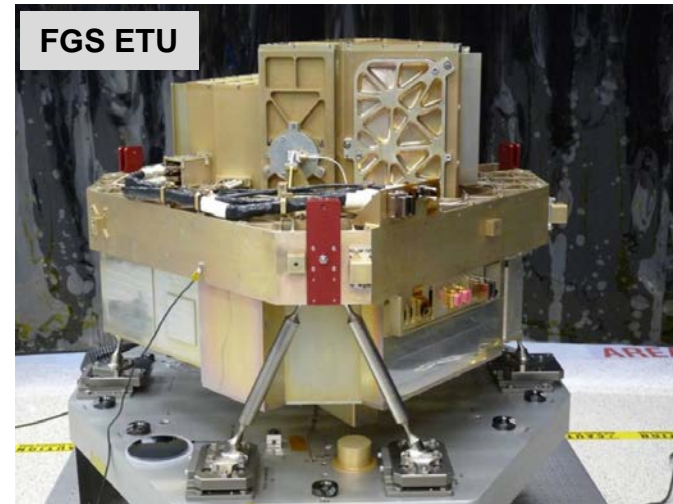
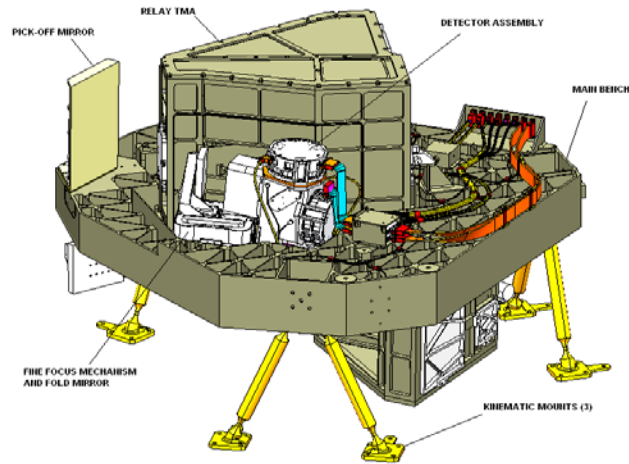
**Flight unit cryo-vacuum testing completed during July 2011**



# MIRI is on schedule for delivery during 2012



# The FGS-Guider and -NIRISS provide telescope pointing control imagery & slitless spectroscopy for Ly- $\alpha$ galaxy surveys and extra-solar planet transits

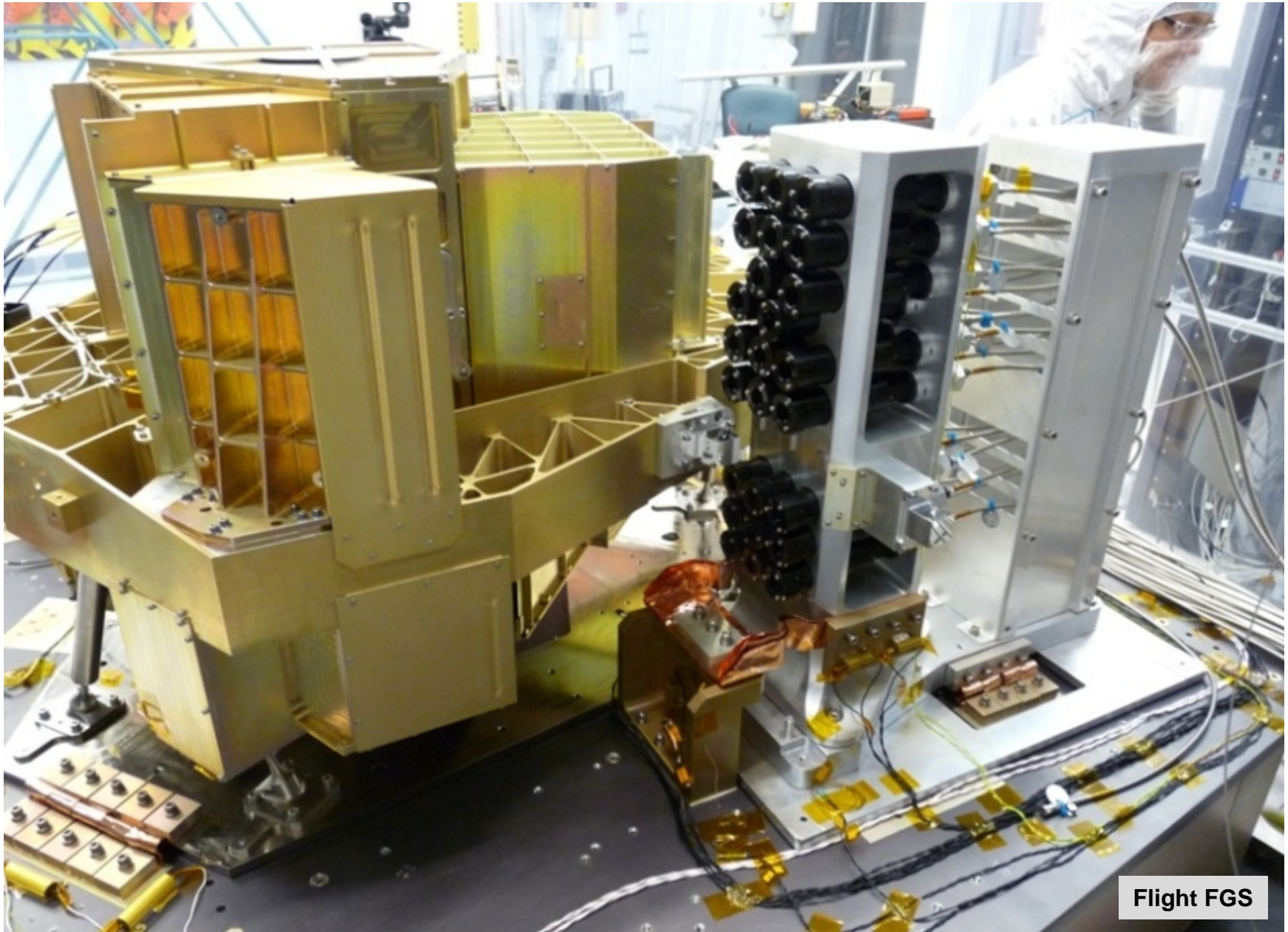


## ■ Developed by the Canadian Space Agency with ComDev

- Broad-band guider (0.6 – 5 microns)
- Field of view: 2.3 x 2.3 arc minutes
- Science imagery:
  - Slitless spectroscopic imagery (grism)
    - R ~ 150, 0.8 – 2.25 microns optimized for Ly alpha galaxy surveys
    - R ~ 700, 0.7 – 2.5 microns optimized for exoplanet transit spectroscopy
  - Sparse aperture interferometric imaging (7 aperture NRM) 3.8, 4.3, and 4.8 microns
- Angular resolution (1 pixel): 68 mas
- Detector type: HgCdTe, 2048 x 2048 pixel format, 3 detectors
- Reflective optics, Aluminum structure and optics

**Flight model cryo-vacuum testing currently underway**

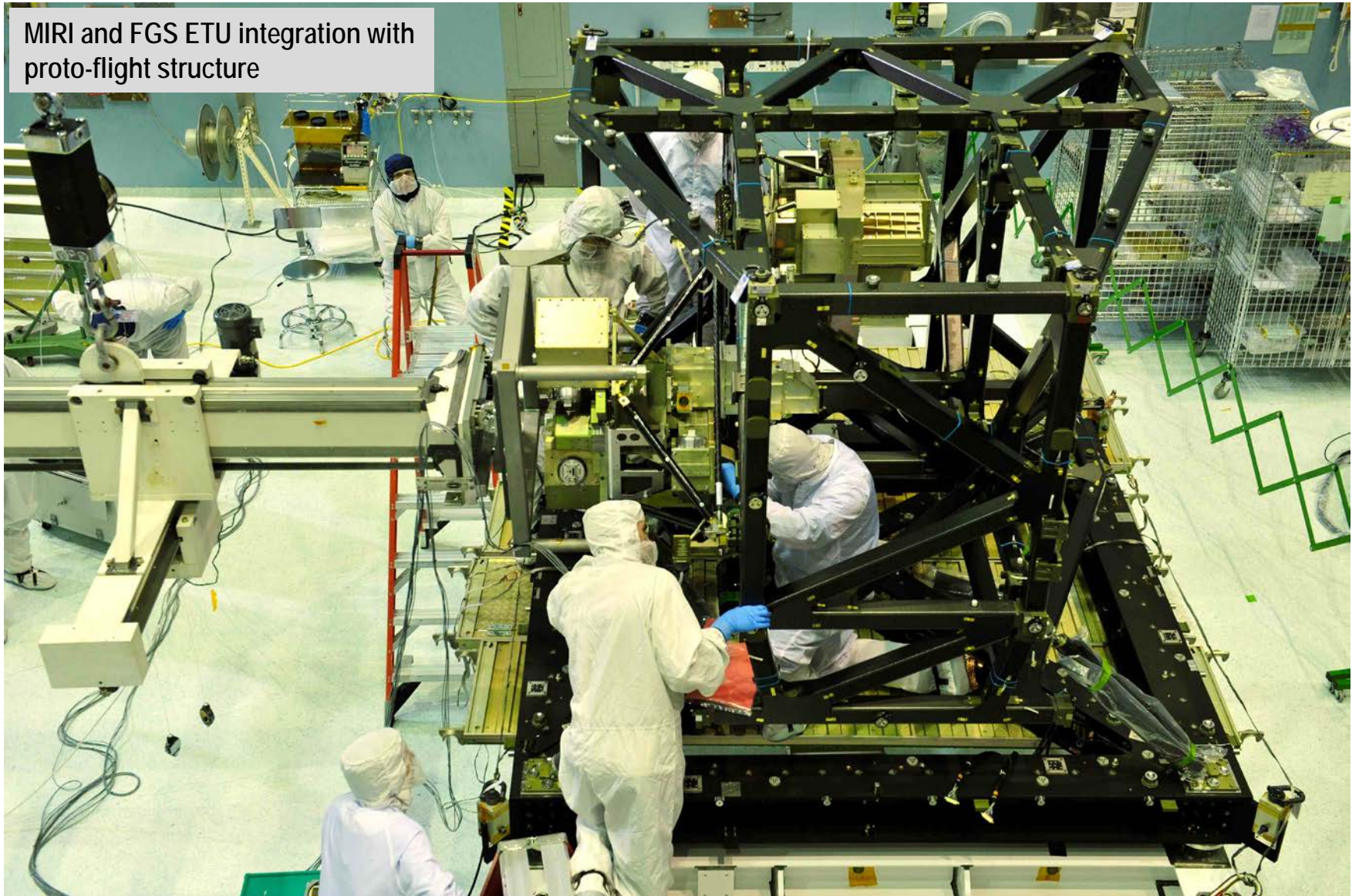
# FGS is on schedule for delivery during 2012



Flight FGS

# Integration of engineering model science instruments with the flight ISIM structure is proceeding well

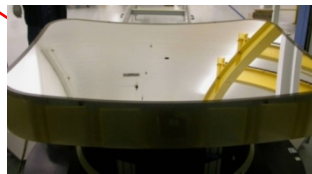
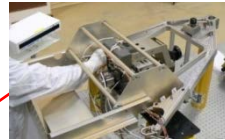
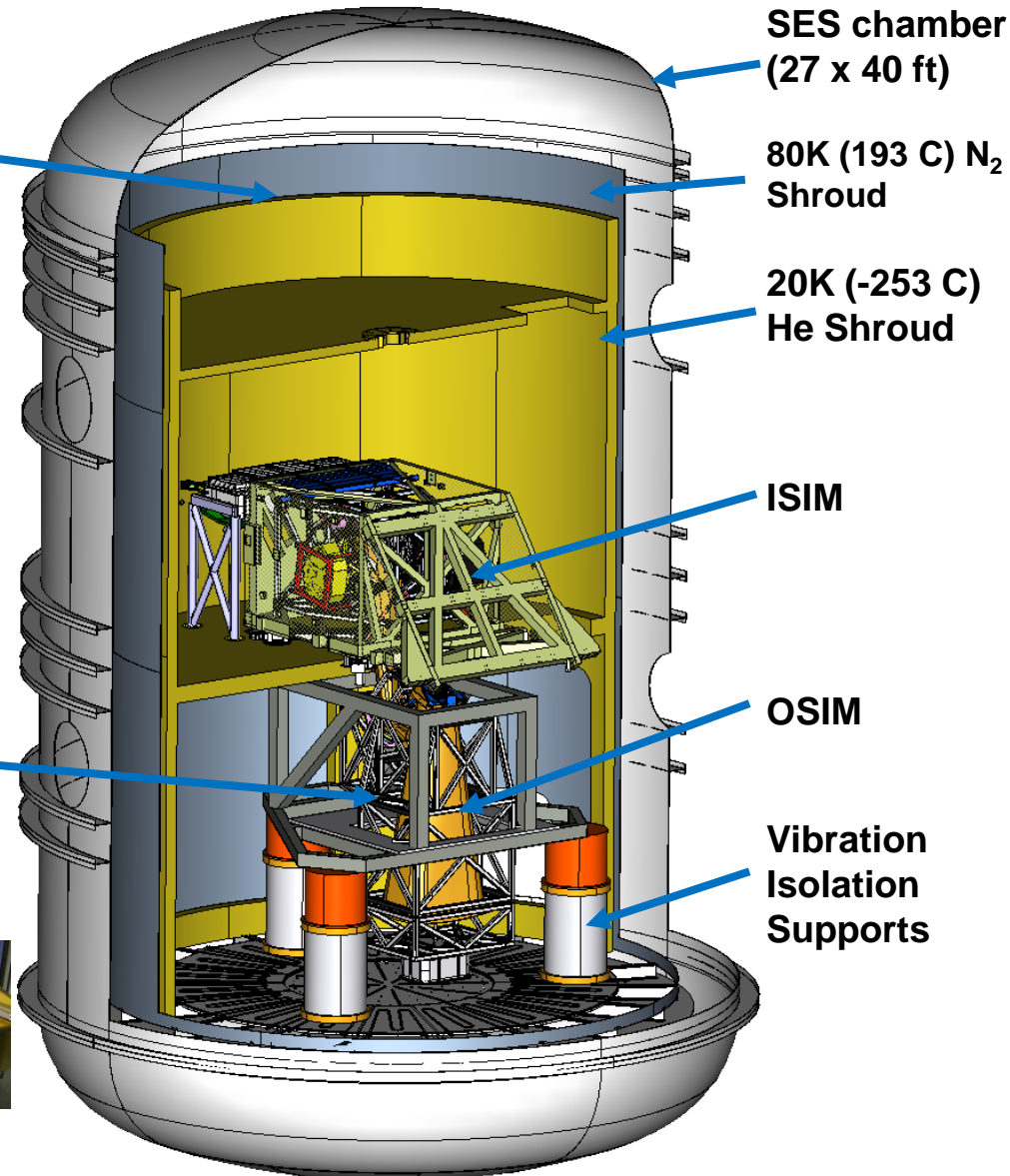
MIRI and FGS ETU integration with proto-flight structure



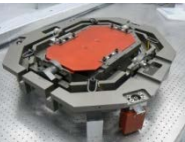
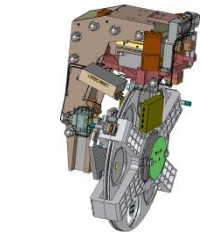
# ISIM will be tested at ~35 K in a space environment simulation chamber using a telescope simulator (OSIM)



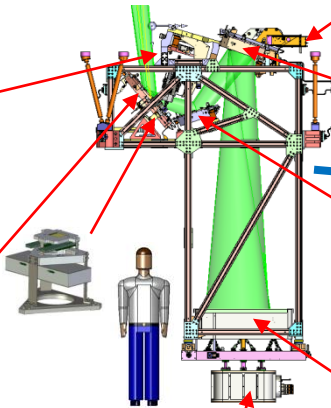
He shroud installation and test completed July 09



OSIM Primary Mirror



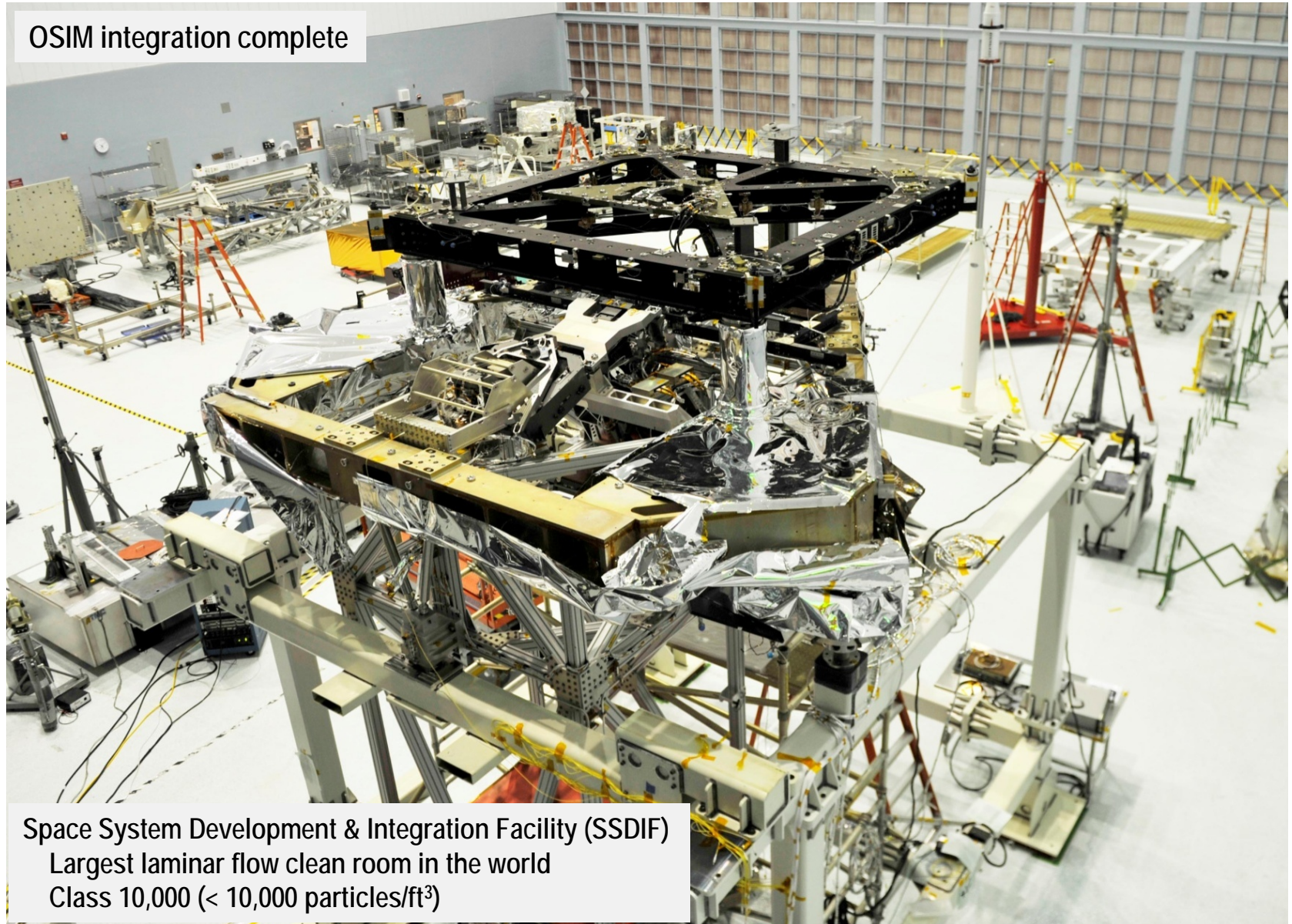
Fold Mirror 3 Tip/Tilt Gimbal Assembly



Alignment Diagnostic Module

# The telescope simulator is on schedule for cryo-vac certification during 2012

OSIM integration complete

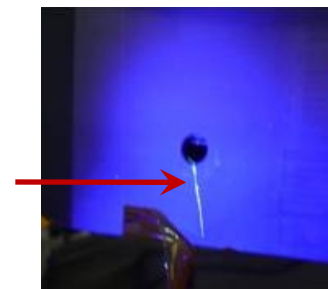


Space System Development & Integration Facility (SSDIF)  
Largest laminar flow clean room in the world  
Class 10,000 ( $< 10,000$  particles/ft<sup>3</sup>)

# ISIM technical issues as of January 2012

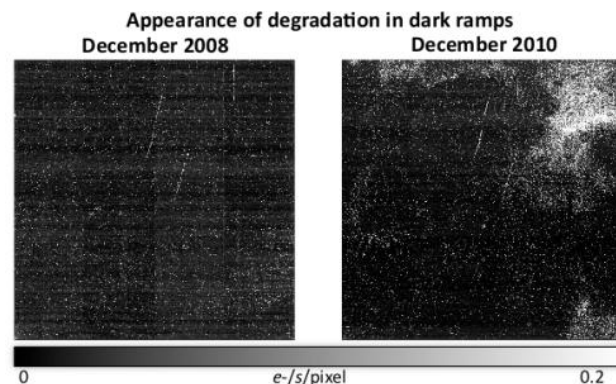
## NIRSpec bench crack:

- Root causes determined; Integration procedure revised to avoid
- Installing flight spare bench
- Puts NIRSpec on ISIM critical path but not expected to delay ISIM delivery

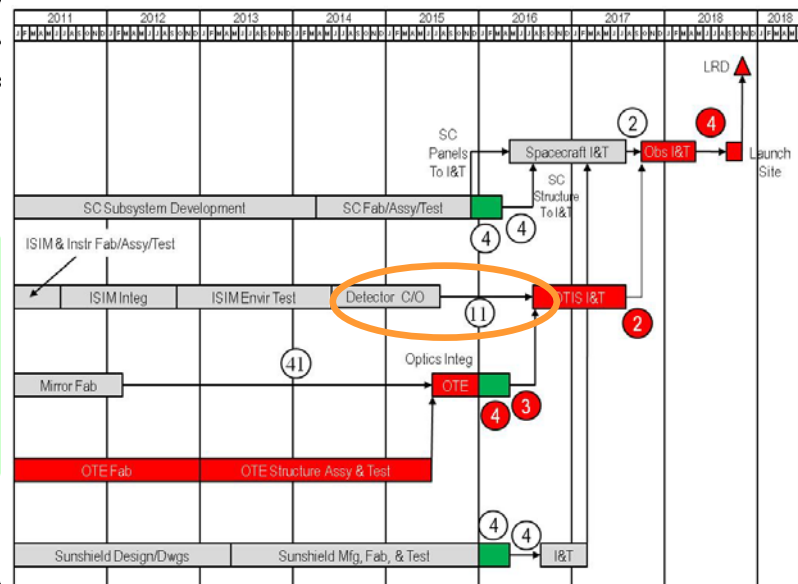


## HgCdTe detector degradation:

- Root cause determined; new design undergoing flight qualification
- Compliance with original performance requirements demonstrated in test
- Qualification program on schedule to complete during June
- Replacement of existing detectors will not delay ISIM delivery



Parameter	Unit	NIRCam Req.	Improved H2RGs	
			H2RG-"A"	H2RG-"B"
Transimpedance gain <sup>1</sup>	$\mu V/e^-$	none	4.46	4.49
QE @ 800 nm	%	none	82	83
QE @ 1000 nm	%	none	89	81
QE @ 1230 nm	%	none	81	78
QE @ 2000 nm <sup>2</sup>	%	$\geq 80\%$	81	81
Median dark current <sup>3</sup>	$e^-/s$	$< 0.01$	0.009	0.004
Median read noise per CDS	$e^-$ rms	$< 21$	9.48	10.07
Median total noise <sup>4</sup>	$e^-$ rms	$< 9$	6.5	5.1
Well capacity	$e^-$	$> 60,000$	97,400	97,800
Crosstalk	%	none	0.8	0.8
Residual image (latency - 2nd read)	%	$< 0.1$	0.07	0.08



**Learn** more at:

[www.jwst.nasa.gov](http://www.jwst.nasa.gov)

[http://webbtelescope.org/webb\\_telescope/progress\\_report/](http://webbtelescope.org/webb_telescope/progress_report/)



**Watch** the JWST being built at:

[www.jwst.nasa.gov/webcam.html](http://www.jwst.nasa.gov/webcam.html)



**Read** about JWST science mission objectives at:

<http://www.jwst.nasa.gov/science.html>

<http://www.stsci.edu/jwst/science/whitepapers/>



**Explore** your science objectives with the JWST exposure time calculator:

<http://jwstetc.stsci.edu/etc/>

**Interact** with the JWST Science Working Group:

<http://www.jwst.nasa.gov/workinggroup.html>

